

98(3): Background Notes on the Noether Theorem

(Lewis H. Ryder, "Quantum Field Theory" (CUP, 1996))

The Noether Theorem is derived from a Lagrangian method described in Ryder's Chapter three. The energy-momentum tensor θ^{μ}_{ν} is defined by Ryder's equation (3.2)

$$\theta^{\mu}_{\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_{\mu}\phi)} \partial_{\nu}\phi - \delta^{\mu}_{\nu} \mathcal{L} \quad (1)$$

It is now assumed that the action S is invariant under a group of transformations x^{μ} and ϕ which for infinitesimal transformations take the form:

$$\Delta x^{\mu} = X^{\mu}_{\nu} \delta\omega^{\nu}, \quad \Delta\phi = \mathbb{F}_{\mu} \delta\omega^{\mu} \quad (2)$$

characterised by the infinitesimal parameter $\delta\omega^{\nu}$. Here X^{μ}_{ν} is a matrix and \mathbb{F}_{μ} is a set of numbers. Here ϕ is a scalar field, δ^{μ}_{ν} is the Kronecker delta, and \mathcal{L} the Lagrangian. If ν is a single index then \mathbb{F}_{μ} is a one-parameter group of transformations. However ν can be a double index, as needed to generate Lorentz transformations. It is now assumed that the action is invariant:

$$\delta S = 0, \quad (3)$$

which implies that:

$$\int_{SR} \left(\frac{\partial \mathcal{L}}{\partial(\partial_{\mu}\phi)} \mathbb{F}_{\nu} - \theta^{\mu}_{\nu} X^{\nu}_{\mu} \right) \delta\omega^{\nu} d\sigma_{\mu} = 0 \quad (4)$$

2) Since S_0^μ is arbitrary:

$$\int_{\partial R} J_\mu^\nu d\sigma_\mu = 0 \quad - (5)$$

where

$$J_\mu^\nu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \overline{F}_\nu - \theta^\mu{}_\nu X^\nu \quad - (6)$$

From Gauss' Theorem:

$$\int_R \partial_\mu J_\mu^\nu d^4x = 0 \quad - (7)$$

and since R is arbitrary:

$$\boxed{\partial_\mu J_\mu^\nu = 0} \quad - (8)$$

The current J_μ^ν is conserved (divergentless), and the ordinary derivative ∂_μ is used because the background spacetime is flat with zero connection. The existence of J_μ^ν follows from the invariance of action under the transformations (2). The conserved or time-independent charge is now defined as:

$$Q_\nu = \int_S J_\mu^\nu d\sigma_\mu \quad - (9)$$

where σ_μ is a space-like hypersurface in Minkowski spacetime. If this surface is defined by:

3)

$$t = \text{constant} \quad (10)$$

then

$$Q_0 = \int_V \dot{T}_0 d^3x \quad (11)$$

where the integral is taken over the volume V . The conservation of Q_0 follows by integrating (8) over

$$\int_V \partial_0 \dot{T}_0 d^3x + \int_V \partial_i \dot{T}_i d^3x = 0 \quad (12)$$

The second term is transformed into a surface integral by the 3-D Gauss theorem and vanishes, so we obtain Noether's Theorem.

$$\boxed{\frac{d}{dt} \int \dot{T}_0 d^3x = \frac{dQ_0}{dt} = 0} \quad (13)$$

Translational Invariance

This is the translation of the origin of space and time, in which case:

$$\Delta x^\mu = \epsilon^\mu, \quad \Delta \phi = 0 \quad (14)$$

e.

$$X^\mu = \delta^\mu, \quad \Phi_\mu = 0 \quad (15)$$

From eqs. (15) and (6):

4) The corresponding conservation law is

$$\frac{d}{dt} \int \theta_{\mu 0} d^3x = 0 \quad - (17)$$

It can be shown that:

$$P_{\mu} = \int \theta_{\mu 0} d^3x \quad - (18)$$

is the 4-momentum or energy-momentum of the scalar field ϕ .

- (19)

Proof

$$\begin{aligned} \int \theta_{\mu 0} d^3x &= \int \left(\frac{\partial \mathcal{L}}{\partial(\partial_0 \phi)} \partial_0 \phi - \mathcal{L} \right) d^3x \\ &= \int \left(\frac{\partial \mathcal{L}}{\partial \dot{\phi}} \dot{\phi} - \mathcal{L} \right) d^3x \quad - (20) \end{aligned}$$

In point particle dynamics the relation between the Hamiltonian and Lagrangian is:

$$H = \sum_i p_i \dot{q}_i - L \quad - (21)$$

and the momentum is defined as:

$$p_i = \frac{\partial L}{\partial \dot{q}_i} \quad - (22)$$

The right hand side of eq. (20) is therefore the energy of the field. Thus $\int \theta_{\mu 0} d^3x$ is a momentum because $\partial \phi / \partial x^{\mu}$ is a 4-vector under Lorentz transformations.

5) The canonical energy-momentum tensor is defined as:

$$D_{\mu} T^{\mu\nu} = 0 \quad - (23)$$

is flat a Minkowski spacetime. In curved spacetime

$$D_{\mu} T^{\mu\nu} = 0 \quad - (24)$$

and the Einstein field equation is derived by:

$$D_{\mu} G^{\mu\nu} = \kappa D_{\mu} T^{\mu\nu} \quad - (25)$$

where $G^{\mu\nu}$ is the Einstein tensor and κ is the Einstein constant. Here:

$$D_{\mu} G^{\mu\nu} = 0 \quad - (26)$$

is the second Bianchi identity.

Rotational Invariance

In this case the action is invariant under spatial rotations:

$$x^i = F^{ij} x^j, \quad F^{ij} = -F^{ji} \quad (i, j = 1, 2, 3) \quad - (27)$$

$$\Delta \phi = 0 \quad - (28)$$

The rotation group is a sub-group of the Lorentz group, so:

$$x^{\mu} = F^{\mu\nu} x^{\nu}, \quad F^{\mu\nu} = -F^{\nu\mu} \quad - (29)$$

This is written as:

$$b) \int x^\mu = X^\mu_{\rho\sigma} \epsilon^{\rho\sigma}, \quad X^\mu_{\rho\sigma} = \frac{1}{2} (\delta^\mu_\rho x_\sigma - \delta^\mu_\sigma x_\rho) \quad - (30)$$

This is the same form as eq. (2) but the index $\rho\sigma$ in eq. (2) has become the double index $\rho\sigma$ of eq. (30). Now use eq. (6) with:

$$\underline{\Phi} = 0, \quad \theta^{\mu\nu} \rightarrow T^{\mu\nu} \quad - (31)$$

to find the three index conserved Noether current:

$$J^{\mu\rho\sigma} = -T^{\mu\kappa} X^{\kappa\rho\sigma} \quad - (31)$$

$$J^{\mu\rho\sigma} = -\frac{1}{2} (T^{\mu\rho} x^\sigma - T^{\mu\sigma} x^\rho) \quad - (32)$$

$$\text{or } \boxed{J^\mu_{\rho\sigma} = -\frac{1}{2} (T^\mu_{\rho\sigma} x^\sigma - T^\mu_{\sigma\rho} x^\rho)} \quad - (33)$$

In Minkowski spacetime:

$$\partial_\mu J^{\mu\rho\sigma} = 0 \quad - (34)$$

and for no-zero connections:

$$\boxed{D_\mu J^{\mu\rho\sigma} = 0} \quad - (35)$$

In analogy with eq. (24) it is seen that $J^{\mu\rho\sigma}$ is the canonical angular momentum

) momentum density tensor. In analogy with the
 Einstein-Hilbert field equation (25) here
 exists a geometrical quantity to which
 $J^{\mu}_{\rho\sigma}$ is proportional. This is the three-
 index Cartan tensor:

$$T^{\mu}_{\rho\sigma} = \Gamma^{\mu}_{\rho\sigma} - \Gamma^{\mu}_{\sigma\rho} \quad - (36)$$

As it follows from (2):

$$J^{\mu}_{\rho\sigma} = \frac{eE^{(0)}}{c} T^{\mu}_{\rho\sigma}, \quad - (37)$$

which is the rotational analogue of the EH field
equation: $G_{\mu\nu} = k T_{\mu\nu}$. $- (38)$

In the quantum theory:

$$eA^{(0)} = \hbar \kappa, \quad - (39)$$

$$E^{(0)} = c b^{(0)} = \kappa c A^{(0)} \quad - (40)$$

$$\text{So } eE^{(0)} = \hbar \kappa^2 c \quad - (41)$$

The Cartan tensor therefore plays the
role of the Einstein tensor $G_{\mu\nu}$ in spatial
rotations as opposed to spatial translations.

8) The most fundamental energy-momentum density is therefore $T_{\mu\nu}$, and the most fundamental angular energy-momentum density is $J^{\mu\nu}$. The four-momentum P_μ is the special case:

$$P_\mu = \int T^\mu_\nu d^3x \quad - (42)$$

and the angular momentum is the special case:

$$J^{\mu\nu} = \int (T^{\mu\alpha} x^\nu - T^{\nu\alpha} x^\mu) d^3x \quad - (43)$$

These special cases are defined by the case:

$$t = \text{constant} \quad - (44)$$

ii) the space-like hypersurface σ_μ used to define the time independent:

$$Q_\nu = \int_\sigma J^\mu_\nu d\sigma_\mu \quad - (45)$$

where:

$$d_\mu J^\mu_\nu = 0. \quad - (46)$$

The Noether theorem is:

$$\frac{dQ_\nu}{dt} = 0, \quad - (47)$$

which the time independence means that momentum/energy

9) and angular momentum energy are conserved, in the sense that they are time-independent. This is because of eq. (44). More generally however energy and momentum can change with time.

Therefore this gives a fundamental justification for defining the electromagnetic field as a rank three canonical tensor.

$$F^{\mu}_{\rho} = \frac{c}{e\omega} J^{\mu}_{\rho} = \frac{E^{(0)}}{\omega} T^{\mu}_{\rho} = A^{(0)} T^{\mu}_{\rho} \quad - (48)$$

where $E^{(0)} = kcA^{(0)} = \omega A^{(0)} \quad - (49)$

In the notation of differential geometry:

$$F^a_{\mu\nu} = A^{(0)} T^a_{\mu\nu} \quad - (50)$$

which is the ECE Ansatz, QED