

# 69(3) : General Theory of the Earth's Precessions (Milankovitch Cycles)

First consider the Lagrangian for the earth ordered as a free symmetric top. This is an idealization because the earth is always influenced by the gravitational fields of the sun, moon and planets.

The Lagrangian is :

$$L = \frac{1}{2} (I_1 \omega_1^2 + I_2 \omega_2^2 + I_3 \omega_3^2) \quad (1)$$

where  $I_1 = I_2 = I_3$  are the principal moments of inertia of the earth, and  $\omega_1, \omega_2$  and  $\omega_3$  are angular velocity components.  $I_1, I_2, I_3$  are the principal moments of inertia of the earth frame (1, 2, 3).

where:

$$\omega_1 = \dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi \quad (3)$$

$$\omega_2 = \dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin \psi \quad (4)$$

$$\omega_3 = \dot{\phi} \cos \theta + \dot{\psi} \quad (5)$$

The Euler Lagrange equations are:

$$\frac{\partial L}{\partial \theta} = \frac{d}{dt} \frac{\partial L}{\partial \dot{\theta}} \quad (6)$$

$$\frac{\partial L}{\partial \phi} = \frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}} \quad (7)$$

$$\frac{\partial L}{\partial \psi} = \frac{d}{dt} \frac{\partial L}{\partial \dot{\psi}} \quad (8)$$

The motion of the free earth is found by solving eqs. (6) to (8) simultaneously for the trajectories  $\vec{r}(t)$ ,  $\phi(t)$  and  $\psi(t)$ . These are the trajectories of the earth in a very weak gravitational field.

When the earth of mass  $m$  is in the gravitational field of an object such as the sun of mass  $M$ , the geometry is as in Fig (1).

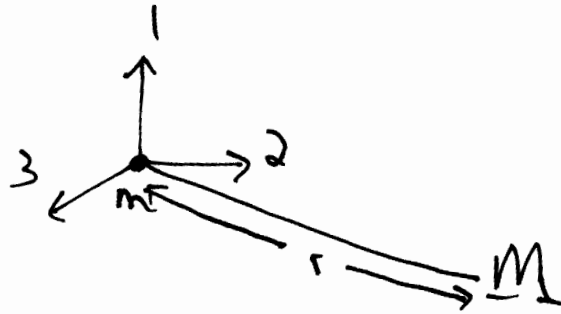


Fig (1)

and the Lagrangian becomes

$$L = \frac{1}{2} m \dot{\vec{r}} \cdot \dot{\vec{r}} + \frac{1}{2} (\bar{I}_1 \omega_1^2 + \bar{I}_2 \omega_2^2 + \bar{I}_3 \omega_3^2) - U \quad (9)$$

where

$$U = -\frac{mM\bar{G}}{r} \quad (10)$$

For self consistency the vector  $\vec{r}$  and the scalar must be transformed into frame (1, 2, 3). In cartesian coordinates:

$$\begin{bmatrix} r_1 \\ r_2 \\ r_3 \end{bmatrix} = \begin{bmatrix} \cos\phi \sin\theta & 0 \\ -\sin\phi \sin\theta & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \cos\theta & 0 & \sin\theta \\ 0 & 1 & 0 \\ -\sin\theta & 0 & \cos\theta \end{bmatrix} \begin{bmatrix} \cos\phi \sin\phi & 0 \\ -\sin\phi \cos\phi & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \end{bmatrix} \quad (11)$$

It is convenient to transform into spherical

'polar coordinates:

$$\begin{aligned} X &= r \sin \theta \cos \phi \\ Y &= r \sin \theta \sin \phi \\ Z &= r \cos \theta \end{aligned} \quad - (12)$$

so:

$$\begin{bmatrix} r_1 \\ r_2 \\ r_3 \end{bmatrix} = r \begin{bmatrix} \cos \phi \sin \theta & \sin \phi \sin \theta & 0 \\ -\sin \phi \sin \theta & \cos \phi \sin \theta & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \cos \theta & 0 & \sin \theta \\ 0 & 1 & 0 \\ -\sin \theta & 0 & \cos \theta \end{bmatrix} \begin{bmatrix} \cos \phi & \sin \phi & 0 \\ -\sin \phi & \cos \phi & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \sin \theta \cos \phi \\ \sin \theta \sin \phi \\ \cos \theta \end{bmatrix} \quad - (13)$$

where

$$r = r(\text{lab}) \quad - (14)$$

meaning that  $r$  is defined in the lab frame. The angles  $\theta$  and  $\phi$  of the spherical polar coordinate system are the same as the Euler angle  $\theta$  and  $\phi$ . Therefore in the  $(1, 2, 3)$  frame:

$$r_1 = r(\text{lab}) f_1(\theta, \phi, \psi) \quad - (15)$$

$$r_2 = r(\text{lab}) f_2(\theta, \phi, \psi)$$

$$r_3 = r(\text{lab}) f_3(\theta, \phi, \psi)$$

and

$$\underline{r}(1, 2, 3) = r_1 \underline{e}_1 + r_2 \underline{e}_2 + r_3 \underline{e}_3 \quad - (16)$$

so

$$r(1, 2, 3) = (f_1^2 + f_2^2 + f_3^2)^{1/2} r(\text{lab}) \quad - (17)$$

Therefore the potential energy in frame  $(1, 2, 3)$  is

$$U = -\frac{mM_6}{r(1, 2, 3)} (f_1^2 + f_2^2 + f_3^2)^{1/2} \quad (18)$$

There is an additional Euler Lagrange equation:

$$\frac{\partial \mathcal{L}}{\partial r(1, 2, 3)} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{r}(1, 2, 3)} \quad (19)$$

The precessions of the Earth in the sun's gravitational field are therefore described by the simultaneous solution of eqs. (6), (7), (8) and (19) to give the trajectories  $r(t)$ ,  $\theta(t)$ ,  $\phi(t)$  and  $\psi(t)$ .

This is the general theory of the precession of the equinox, Milankovitch cycles, and similar.

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