

1 Notes 124(1): Spiral Galaxy

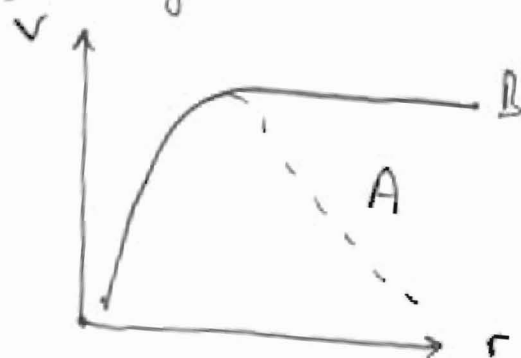
The basic equation is:

$$\underline{\nabla} \cdot \underline{g} = 4\pi G (R - \omega T) \quad - (1)$$

where:

$$\underline{g} = c k \underline{J} \quad - (2)$$

where \underline{J} is an angular momentum. The velocity curve of the galaxy is:



The Newtonian region A is described by:

$$\underline{\nabla} \cdot \underline{g} = 4\pi G R = 4\pi G \rho \quad - (3)$$

i.e.

$$\underline{\nabla} \cdot \underline{J} = 4\pi c k G R \quad - (4)$$

In order to describe the region B, we use

$$R = \omega T \quad - (5)$$

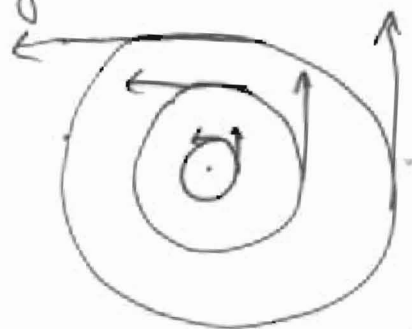
so

$$\underline{\nabla} \cdot \underline{g} = \underline{\nabla} \cdot \underline{J} = 0 \quad - (6)$$

Therefore:

$$\underline{J} = -J_y \underline{i} + J_x \underline{j} \quad - (7)$$

and look follows:



This is beginning to look like a spiral galaxy
a whirlpool.

If the angular momentum is constant:

$$|\underline{J}| = (J_x^2 + J_y^2)^{1/2} = \text{constant} \quad - (8)$$

i.e.

$$J = mrv = mr^2\omega = \text{constant} \quad - (9)$$

and

$$v = \omega r \quad - (10)$$

where

$$\omega = \frac{d\theta}{dt} \quad - (11)$$

If
the

$$\omega r = \text{constant} \quad - (12)$$

$$\boxed{v = \text{constant}} \quad - (13)$$

is a regular B.

We have:

$$J = mrv \quad - (14)$$

3) and if v is constant, J is proportional to r .
The angular velocity is:

$$\frac{d\theta}{dt} = \frac{mv}{r} = \frac{\text{constant}}{r} \quad - (15)$$

So: $\theta = \frac{\text{constant}}{r} \int_0^{\tau} dt \quad - (16)$

$$\theta = \frac{\text{constant} \cdot \tau}{r} \quad - (17)$$

which is a spiral as in paper 76.

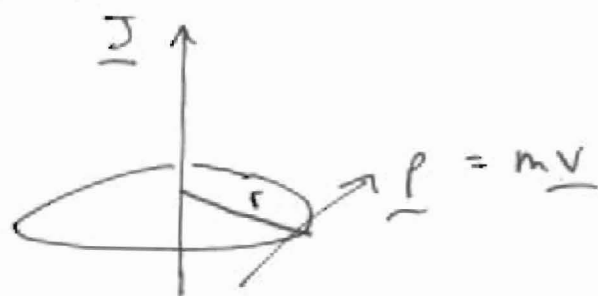
124(2): Angular Momentum is a Constant of Motion

Consider a particle of mass m moving in a central force field described by the potential U . If the latter depends only on the distance of the particle from the force centre and not on orientation, the angular momentum of the system is conserved:

$$\underline{J} = \underline{r} \times \underline{p} = \text{constant} \quad - (1)$$

(S.D. Maria and S.D. Thakur, "Classical Dynamics", HCB 1988, 3rd ed., page 246).

The Lagrangian analysis of the problem defines the Lagrangian in plane polar coordinates r and θ as follows:



$$L = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2) - U(r) \quad - (2)$$

The angular momentum conjugate to the coordinate θ is conserved:

$$\dot{p}_\theta = \frac{\partial L}{\partial \theta} = 0 = \frac{d}{dt} \frac{\partial L}{\partial \dot{\theta}} \quad - (3)$$

i.e.

$$J = p_\theta = \frac{\partial L}{\partial \dot{\theta}} = m r^2 \dot{\theta} = \text{constant} \quad - (4)$$

The angular velocity is defined by:

$$\omega = \dot{\theta} = \frac{d\theta}{dt} \quad - (5)$$

so

$$\boxed{J = m r^2 \omega = m r v = \text{constant}} \quad - (6)$$

The angular momentum \underline{J} is the first integral of the motion. The infinitesimal area dA is:

$$dA = \frac{1}{2} r^2 d\theta \quad - (7)$$

and the areal velocity is:

$$\frac{dA}{dt} = \frac{1}{2} r^2 \frac{d\theta}{dt} = \frac{1}{2} r^2 \omega \quad - (8)$$

$$\boxed{\frac{dA}{dt} = \frac{J}{2m} = \text{constant}} \quad - (9)$$

This is Kepler's second law of planetary motion.

This is true for all central-force motion, not just for an inverse square law.

If the plane of, for example, a galaxy is in $X-Y$, the angular momentum is:

$$\underline{J} = m r^2 \omega \underline{k} \quad - (10)$$

where:

$$\underline{r} = r_x \underline{i} + r_y \underline{j}, \quad - (11)$$

$$r = |\underline{r}| = (r_x^2 + r_y^2)^{1/2}, \quad - (12)$$

$$\text{so } \underline{\nabla} \cdot \underline{J} = \frac{\partial J}{\partial z} = 0 \quad - (13)$$

$$\text{and } \underline{p} = p_x \underline{i} + p_y \underline{j}. \quad - (14)$$

The ECE equation of motion is:

$$\underline{\nabla} \cdot \underline{g} = c^2 (R - \omega T) \quad - (15)$$

3) where:

$$\underline{g} = c^2 \left(T^{010} \underline{i} + T^{020} \underline{j} + T^{030} \underline{k} \right) \quad - (16)$$

Now integrate over volume:

$$T^{i0} = \int T^{0i0} dV = V T^{0i0} \quad - (17)$$

$$i = 1, 2, 3.$$

and define the angular momentum tensor:

$$\underline{J}^{i0} = \frac{c}{k} T^{i0} \quad - (18)$$

Therefore:

$$\underline{g} = \frac{ck}{V} \underline{J} \quad - (19)$$

Units Check

$$g = \text{m s}^{-2}; \quad c = \text{m s}^{-1}, \quad k = \text{m kg m}^{-1}, \quad \checkmark \checkmark$$

$$J = \text{kg m}^2 \text{ s}^{-1}, \quad V = \text{m}^3$$

Therefore:

$$\underline{\nabla} \cdot \underline{g} = \frac{ck}{V} \underline{\nabla} \cdot \underline{J} = 4\pi G \rho \quad - (20)$$

$$\text{where: } G = \frac{c^2 k}{8\pi} \quad - (21)$$

Therefore

$$\underline{\nabla} \cdot \underline{J} = \rho = \frac{1}{2} V c \rho \quad - (22)$$

4) and:

$$\underline{\nabla} \cdot \underline{J} = \frac{\nabla c}{k} (R - \omega T) \quad - (23)$$

If it is assumed that the mass density is

$$\rho = \frac{m}{V} \quad - (24)$$

then:

$$\underline{\nabla} \cdot \underline{J} = \frac{1}{2} mc \quad - (25)$$

Denote:

$$\nabla (R - \omega T) := (R - \omega T)' \quad - (26)$$

then:

$$\underline{\nabla} \cdot \underline{J} = \frac{c}{k} (R - \omega T)' \quad - (27)$$
$$= \frac{1}{2} mc$$

and

$$m = \frac{2}{k} (R - \omega T)' \quad - (28)$$

Therefore mass is proportional to $(R - \omega T)'$.
The fundamental dynamics of the Newton and Coulomb laws are given by eq. (25), where mass m acts as a source for the field \underline{J} .

5) Central Motion

In central motion:

$$\underline{\nabla} \cdot \underline{J} = 0 \quad - (29)$$

and so:

$$\boxed{R = \omega T} \quad - (30)$$

is a general law of central motion. Therefore Kepler's second law is eq. (30). Eq. (30) near
 that:

$$m = 0 \quad - (31)$$

meaning that there is no mass outside the plane of the orbit.

Kepler's First and Third Laws

These depend specifically on the inverse square law of Newton. If Newtonian dynamics are defined by:

$$\omega \rightarrow 0 \quad - (32)$$

(specifically a gas to zero), then the first and third laws are given by:

$$\underline{\nabla} \cdot \underline{J} = \frac{c}{k} R \quad - (33)$$

$$\boxed{R' = R^{10} + R^{20} + R^{30}} \quad - (34)$$

124(3) : Relation Between Torque and Spi (Contra) for General Central orbits.

The general relation between force law and orbit is given from a Lagrangian analysis by eq. (7.20) of Maria and Thornton. This can be expressed as:

$$F(r) = -m v^2 \left(\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right) \quad (1)$$

where m is the mass of the attracting particle. The potential energy is:

$$U(r) = - \int F dr = m \int v^2 \left(\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right) dr \quad (2)$$

and the gravitational potential is:

$$\Phi = \frac{1}{m} U(r) = \int v^2 \left(\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right) dr \quad (3)$$

Maria and Thornton write eq. (1) as:

$$\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} = - \frac{m r^2}{J^2} F(r) \quad (4)$$

where $J = m r^2 \omega = m r^2 \dot{\theta} = \text{constant}$.

Therefore:

(5)

$$F(r) = -r^2 \omega^2 m \left(\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right) \quad - (6)$$

Define:

$$v = r\omega \quad - (7)$$

to be the orbital linear velocity. The angular velocity is:

$$\omega = \frac{d\theta}{dt} \quad - (8)$$

In ECE theory: - (9)

$$\underline{g} = -\underline{\nabla} \underline{\Phi} + \underline{\omega} \underline{\Phi} = c^2 \underline{T}$$

Eq. (9) gives a relation between the torsion \underline{T} and the spin convention $\underline{\omega}$ for a given $\underline{\Phi}$, defined by eq. (3). From Kepler's second law:

$$\frac{dA}{dt} = \frac{1}{2} r^2 \omega = \text{constant} \quad - (10)$$

Therefore:

$$v = \frac{2}{r} \frac{dA}{dt} \quad - (11)$$

Therefore

$$v = \frac{k}{r} \quad - (12)$$

where:

$$k = 2 \frac{dA}{dt} = \text{constant} \quad - (13)$$

for all central orbits

Therefore in eq. (3):

$$\underline{\Phi} = \int \frac{k^2}{r^2} \left(\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right) dr \quad - (14)$$

where: $k = 2 \frac{dA}{dt} \quad - (15)$

Since k is a constant:

$$\underline{\Phi} = k^2 \int \left(\frac{1}{r^2} \left(\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right) dr \right) \quad - (16)$$

Newtonian orbits

These are given by:

$$\frac{1}{r} = \frac{1}{d} (1 + \epsilon \cos \theta) \quad - (17)$$

(M&T eq. (7.41)), and by:

$$\underline{\Phi} = - \frac{Gm}{r} \quad - (18)$$

(M&T eq. (4.6)). The Newtonian orbits are particular cases of eq. (16)

Log Spiral orbits (M&T page 250)

These are: $r = k \exp(d\theta) \quad - (19)$

and $\underline{\Phi} = k^2 (1+d^2) \int \frac{1}{r^2} dr$

$$\underline{\Phi} = - \frac{1}{2} k^2 (1+d^2) \frac{1}{r^2} \quad - (20)$$

24 (4): Kepler's Equation

In astronomy we need the function $\theta(t)$, so the direction of an orbiting object may be found at any time. Planetary motion for example is described by:

$$\frac{d}{r} = 1 + e \cos \theta \quad - (1)$$

i.e. the equation of a conic section with one focus at the origin (Moria & Thornton eq. (7.41)). Here e is the eccentricity and $2a$ is the latus rectum.

Define the period of the orbit as τ . This is the time taken for the radius vector to sweep out the entire area πab of an elliptical orbit. By Kepler's

second law, the area $(\pi ab / \tau)t$ is swept out (M & T page 261). Thus:

$$\int dA = \frac{\pi ab}{\tau} t \quad - (2)$$

where

$$dA = \frac{1}{2} r^2 d\theta \quad - (3)$$

If $\theta = 0$ at $t = 0$:

$$\frac{\pi ab}{\tau} t = \frac{1}{2} \int_0^\theta r^2 d\theta \quad - (4)$$

where, from eq. (1):

$$r = \frac{d}{1 + e \cos \theta} \quad - (5)$$

so

$$\frac{\pi ab}{\tau} t = \frac{d^2}{2} \int_0^\theta \frac{d\theta}{(1 + e \cos \theta)^2} \quad - (6)$$

Now we:

$$ab = a^2 (1 - e^2)^{-3/2} \quad - (7)$$

So:

$$\frac{2\pi t}{\tau} = 2 \tan^{-1} \left(\left(\frac{1-e}{1+e} \right)^{1/2} \tan \frac{\theta}{2} - \frac{e (1-e^2)^{1/2} \sin \theta}{1+e \cos \theta} \right) \quad - (8)$$

(M & T eq. (7.53)).

This equation must be inverted to give θ as a function of t . It may be possible to use Maxima to do this. An approximate result is:

$$\begin{aligned} \theta(t) = & \frac{2\pi t}{\tau} + 2e \sin \left(\frac{2\pi t}{\tau} \right) + \frac{5}{4} e^2 \sin \left(\frac{4\pi t}{\tau} \right) \\ & + \frac{1}{12} e^3 \left(13 \sin \left(\frac{6\pi t}{\tau} \right) - 3 \sin \left(\frac{2\pi t}{\tau} \right) \right) \\ & + \dots \end{aligned} \quad - (9)$$

This problem was also solved by Kepler's Equation:

$$\frac{2\pi t}{\tau} = \phi - e \sin \phi \quad - (10)$$

where:

$$\tan \frac{\theta}{2} = \left(\frac{1+\epsilon}{1-\epsilon} \right)^{1/2} \tan \frac{\phi}{2} \quad - (11)$$

So ϕ is found as a function of t from eq. (10), by inversion, and ϕ is related to θ by eq. (11).

Velocity as a Function of the Radius Vector
This is found from Kepler's equation using:

$$v^2 = \dot{x}^2 + \dot{y}^2 \quad - (12)$$

$$= a^2 \dot{\phi}^2 (1 - \epsilon^2 \cos^2 \phi) \quad - (13)$$

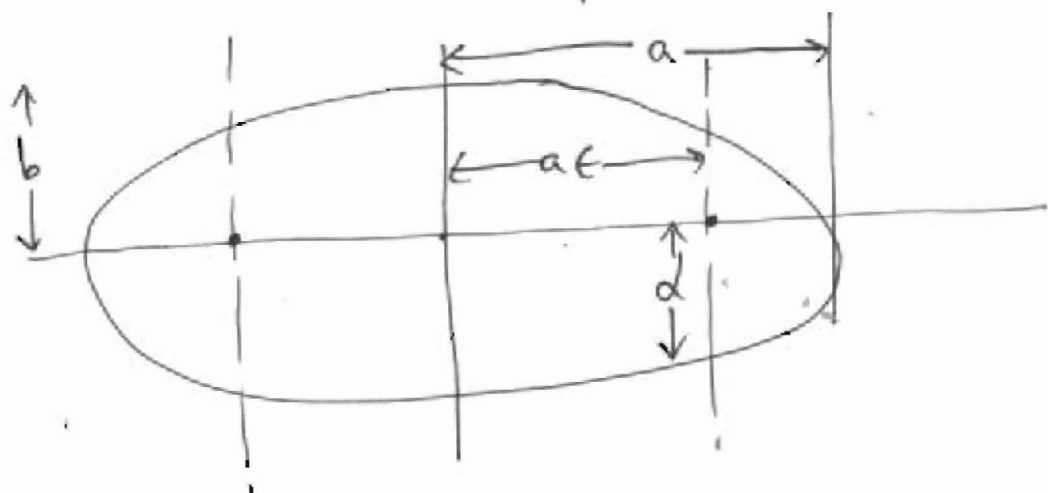
(in & T eq. (7.68)). So:

$$v^2 = \frac{k}{\mu} \left(\frac{2}{r} - \frac{1}{a} \right) \quad - (14)$$

Kepler's third law is:

$$\tau^2 = \frac{4\pi^2 \mu}{k} a^3 \quad - (15)$$

The inverse sq. law is: $F = -k/r^2$
 $k = m_1 m_2 G$



Invaria Method for Any Central Force Law
The angular velocity ω is defined by:

$$\omega = \frac{d\theta}{dt} = \frac{J}{mr^2} = \frac{\text{constant}}{r^2} \quad (1)$$

so:

$$\theta = \text{constant} \int \frac{1}{r^2} dt \quad (2)$$

In their chapter seven, Maria and Thantia give:

$$\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} = -\frac{mr^2}{J^2} F(r) \quad (3)$$

which is valid for any central force law $F(r)$. This gives plenty of scope for animations.

From eq. (3):

$$\frac{d\theta}{dt} = \frac{J}{mr^2} = -JF(r) \left(\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right)^{-1} \quad (4)$$

where

$$J = mr^2 \omega = \text{constant of motion} \quad (5)$$

So the orbit for any central force law is:

$$\theta = -J \int \frac{1}{F(r) \left(\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right)^{-1}} dt \quad (6)$$

The Inverse Square Force Law

This is: $F(r) = -\frac{k}{r^2} = -\frac{m_1 m_2 G}{r^2} \quad (7)$

2) Eq. (7) gives:

$$\frac{1}{r} = \frac{1}{d} (1 + e \cos \theta) \quad - (8)$$

where d and e are the half latus rectum and eccentricity of an ellipse. So the dependence of $1/r$ on θ is given by eq. (8) for inverse square law.

If the orbit is a circle:

$$e = 0 \quad - (9)$$

so:

$$\theta = - \int r F(r) dt \quad - (10)$$

$$= Jk \int \frac{1}{r} dt$$

$$\theta = \frac{Jkt}{r} \quad - (11)$$

because r is constant in time for a circular orbit.

More generally, r is a function of t in eq. (6). This can be seen from the fact that θ is a function of t in eq. (8). The dependence of r on t for the inverse square law is given by Kepler's equation:

$$v^2 = \frac{k}{m} \left(\frac{2}{r} - \frac{1}{a} \right) \quad - (12)$$

where

$$\underline{v} = \frac{dx}{dt} \underline{i} + \frac{dy}{dt} \underline{j} \quad - (13)$$

$$v^2 = x^2 + y^2 \quad - (14)$$

So:

$$\frac{dr}{dt} = \left(\frac{k}{m} \left(\frac{2}{r} - \frac{1}{a} \right) \right)^{1/2}$$

$$\frac{dt}{dr} = \left(\frac{k}{m} \left(\frac{2}{r} - \frac{1}{a} \right) \right)^{-1/2} \quad - (15)$$

$$t = \int \left(\frac{k}{m} \left(\frac{2}{r} - \frac{1}{a} \right) \right)^{-1/2} dr \quad - (16)$$

1) The dependence of t on r may be animated from eq. (16). From eqs. (6) and (15):

$$\theta = - \int \sqrt{F(r) \left(\frac{d^2}{dt^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right)^{-1} \left(\frac{k}{m} \left(\frac{2}{r} - \frac{1}{a} \right) \right)^{-1/2}} dr \quad - (17)$$

2) and the dependence of θ on r may be animated from this equation.

3) The dependence of r on t may be found by inverting the results of eq. (16) and parametrizing numerically.

4) The simplest method is to use eq. (2) and eq. (16). The latter is integrated by Maxima to give t as a function of r , and the result inverted to give r as a function of t .

4) Log Spiral Orbit

This is worked out completely by M & T, pp. 250 and 251. If a particle moves on a log spiral:

$$r = k \exp(d\theta) \quad - (1)$$

where k and d are constants. Eq. (1) gives:


$$F(r) = - \frac{J^2}{mr^3} (1+d^2) \quad - (2)$$

and
$$\theta = \frac{1}{2d} \log \left(\frac{2dJt}{mk^2} + c \right) \quad - (3)$$

with
$$r = \left(\frac{2dJt}{m} + k^2c \right)^{1/2} \quad - (4)$$

Animate eqs. (1), (3) and (4).

These were first discussed by Roger Cotes (1682 - 1716) and the orbits are Cotes' Spirals.



4 (b): Effect of Constant Specific Torion on the Newtonian Attraction of Stars in a Galaxy

In general, the total energy in a central orbit of any kind is constant:

$$E = T + U = \text{constant} \quad - (1)$$

$$E = \frac{1}{2} m \dot{r}^2 + \frac{1}{2} \frac{J^2}{m r^2} + U(r) \quad - (2)$$

Therefore:

$$\frac{dr}{dt} = \left(\frac{2}{m} (E - U) - \frac{J^2}{m^2 r^2} \right)^{1/2} \quad - (3)$$

where E and J are constants. The inverse square potential is:

$$U = m \Phi = - \frac{G m^2}{r} \quad - (4)$$

Therefore, as in M & T chapter 7:

$$\theta = \int \frac{(J/r^2) dr}{2m \left(E - U(r) - \frac{J^2}{2mr^2} \right)^{1/2}} \quad - (5)$$

One may animate θ as a function of r .

The effective potential is:

$$V(r) = U(r) + \frac{J^2}{2mr^2} \quad - (6)$$

$$V(r) = - \frac{Gm^2}{r} + \frac{J^2}{2mr^2}$$

2) and consists of the attraction and centrifugal repulsion, i.e. respectively negative and positive.

The Effect of Constant Spacetime Torsion

From previous work it is known that torsion, when integrated over volume, is proportional to angular momentum. Therefore in a galaxy, there is an additional repulsive potential due to constant spacetime torsion:

$$\boxed{V(\text{torsion}) = \frac{J_T^2}{2mr^2}} \quad - (7)$$

The extra repulsive force is:

$$F(\text{torsion}) = - \frac{\partial V(\text{torsion})}{\partial r} = + \frac{J_T^2}{mr^3} \quad - (8)$$

From problem (7.22) of m & T , if:

$$J_0^2 := J_T^2 (1 + \alpha_0) \quad - (9)$$

then the orbit produced by eq. (8) is a logarithmic spiral:

$$\boxed{r = k \exp(\alpha_0 \theta)} \quad - (10)$$

Therefore this simple model produces the main features of a spiral galaxy. The Newtonian curve of θ versus r is given by eq. (5), and the spiral orbits by eq. (10):

$$\theta = \frac{1}{d_0} \log_e \frac{r}{R} \quad - (11)$$

Eq. (5) gives:

$$\cos \theta = \frac{1}{F} \left(\frac{d}{r} - 1 \right)$$

$$\theta = \cos^{-1} \left(\frac{1}{F} \left(\frac{d}{r} - 1 \right) \right) \quad - (12)$$

Animations can be made of θ as a function of r for eqns. (11) and (12).

The arms of a whirlpool galaxy are logarithmic spirals. We have:

$$\frac{dr}{d\theta} \frac{d\theta}{dt} = br \quad - (13)$$

$$\text{and } v = \frac{dr}{d\theta} \frac{d\theta}{dt} = \omega \frac{dr}{d\theta} = b\omega r \\ = \text{constant} \quad - (14)$$

1) Note 124(8): Logarithmic spiral orbit

Consider the Lagrangian: $L = \frac{1}{2} m (\dot{r}^2 - r^2 \dot{\theta}^2) - (1)$

and the Euler Lagrange equations: $\frac{\partial L}{\partial r} = \frac{d}{dt} \frac{\partial L}{\partial \dot{r}} - (2)$

Thus: $\frac{\partial L}{\partial r} = m r, \quad \frac{d}{dt} \frac{\partial L}{\partial \dot{r}} = m \ddot{r} - (3)$

$\frac{\partial L}{\partial r} = -m r \dot{\theta}^2 - (4)$

Thus: $\ddot{r} = -r \dot{\theta}^2 - (5)$

Making the transform of variable: $u = \frac{1}{r} - (6)$

and using the method of Maria and Thoma page

249: $\ddot{r} = -\frac{J^2}{m^2} u^2 \frac{d^2 u}{d\theta^2} - (7)$

$r \dot{\theta}^2 = \frac{J^2}{m^2} u^3 - (8)$

The centrifugal force is:

Using eq. (8):

$$F = -\frac{dU}{dr} = m r \dot{\theta}^2 = \frac{J^2}{m r^3} \quad (9)$$

which is the centrifugal force outwards.

Therefore

$$m r \ddot{r} - \frac{dU}{dr} = 0 \quad (10)$$

and the orbital equation is

$$\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) = -\frac{1}{r} \quad (11)$$

This is satisfied by $\log \text{ spiral}$

$$r = e^{\theta} \quad (12)$$

We have:

$$r(t) = \left(\frac{2Jt}{m} + C \right)^{1/2} \quad (13)$$

$$\theta(t) = \frac{1}{2} \log \left(\frac{2Jt}{m} + C \right) \quad (14)$$

giving the evolution of a whirlpool galaxy.

3)

The velocity of a star is:

$$v^2 = \dot{r}^2 + r^2 \dot{\theta}^2 \quad (15)$$

and may be calculated from eqs. (13) and (14).

Animations

Eqns. (13) to (15) may be animated. If

the velocity v becomes constant, then:

$$\dot{r}^2 + r^2 \dot{\theta}^2 = \text{constant} \quad (16)$$

and it is seen that eqs. (13), (14) and (16) give an equation t . At some critical time the galaxy has evolved to a constant v , and the arms of the spiral become straightened out & observed

124(9): Velocity Curve of a Log Spiral Orbit.

The velocity of the star in a logarithmic spiral orbit is:

$$v^2 = \dot{r}^2 + r^2 \dot{\theta}^2 \quad - (1)$$

where:

$$\dot{\theta} = \frac{J}{m r_0^2} e^{-2\theta} \quad - (2)$$

(Mera and Thonka, eq. (7.23), page 250),

and:

$$r = \frac{d_0 r_0 J}{m r^2} e^{d\theta} \quad - (3)$$

(Mera and Thonka, eq. (7.28), page 251).

The log spiral is:

$$r = r_0 e^{d\theta} \quad - (4)$$

so

$$\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) = \frac{d^2}{r} \quad - (5)$$

Therefore:

$$v^2 = \left(\frac{d_0 r_0 J}{m} \right)^2 \frac{e^{2d\theta}}{r^4} + \left(\frac{J}{m r_0^2} \right)^2 r^2 e^{-4d\theta} \quad - (6)$$

2) Define:

$$A = \frac{d \cdot r_0 J}{m}, \quad B = \frac{J}{m r_0^2} \quad - (7)$$

Req:

$$v^2 = \left(\frac{A e^{d\theta}}{r^2} \right)^2 + \left(B r e^{-2d\theta} \right)^2 \quad - (8)$$

Limits

1) As $r \rightarrow \infty$ - (9)

$$v \rightarrow B r e^{-d\theta} \quad - (10)$$

2) As $r \rightarrow 0$ - (11)

$$v \rightarrow \infty \quad - (12)$$

In R limit (9):

$$v \rightarrow \frac{J r}{m r_0^2} e^{-d\theta} \quad - (13)$$

Using eq. (4):

$$v \rightarrow \frac{J}{m r_0} = \text{constant}$$

for any r. The graph of v against r is a plateau as observed experimentally.

1) 124(10): The Potential Energy Generated by
Spacetime Torsion.

Consider a star moving along the radial direction \underline{e}_r with velocity:

$$\underline{v}_r = \dot{r} \underline{e}_r \quad - (1)$$

By Newton's first law it would move along this direction permanently unless acted upon by an external force, \underline{F} . This force produces work done on the star:

$$W_{12} = \int_1^2 \underline{F} \cdot d\underline{r}, \quad - (2)$$

work which transforms the star from condition 1 to 2. The kinetic energy generated by \underline{v} of eq. (1) is:

$$T = \frac{1}{2} m \dot{r}^2. \quad - (3)$$

The torsion of spacetime does work on the star while keeping T constant. The potential energy transferred to the star from spacetime is defined by:

$$\int_1^2 \underline{F} \cdot d\underline{r} = U_1 - U_2 \quad - (4)$$

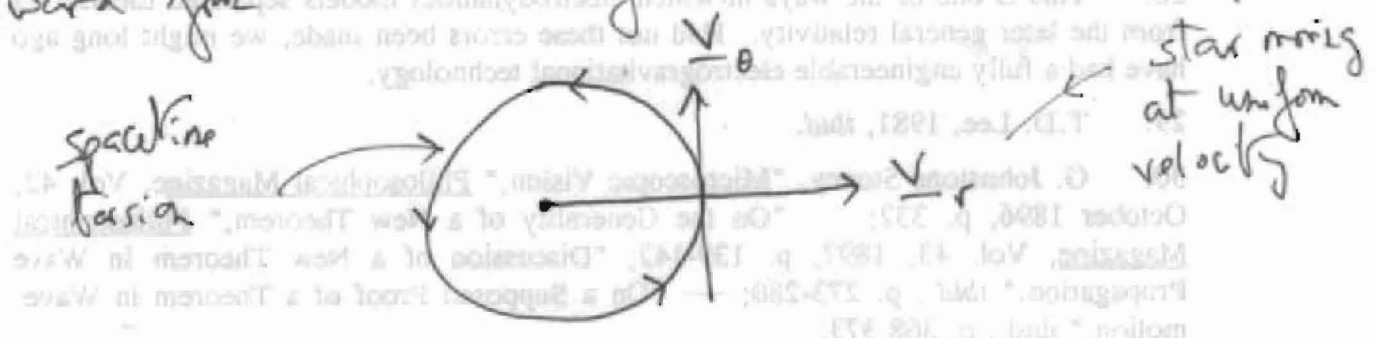
so:

$$\underline{F} = -\underline{\nabla}U \quad - (5)$$

This potential energy moves the star in a direction transverse to \underline{e}_r , producing the

2) velocity:
$$\underline{v}_\theta = r \dot{\theta} \underline{e}_\theta \quad (6)$$

This process may be thought of as a particle moving outwards from the centre of a rotating platform:



The total velocity is
$$\underline{v} = \dot{r} \underline{e}_r + r \dot{\theta} \underline{e}_\theta \quad (7)$$

and the total energy is
$$E = \frac{1}{2} m \underline{v} \cdot \underline{v} \quad (8)$$

$$E = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2)$$

where
$$E = T + U \quad (9)$$

where
$$T = \frac{1}{2} m \dot{r}^2 \quad (10)$$

$$U = \frac{1}{2} m r^2 \dot{\theta}^2 \quad (11)$$

Re Lagrangian is
$$\mathcal{L} = T - U = \frac{1}{2} m (\dot{r}^2 - r^2 \dot{\theta}^2) \quad (12)$$

3) The Euler Lagrange equation is:

$$\frac{\partial \mathcal{L}}{\partial r} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{r}} \quad (13)$$

and

$$\frac{\partial \mathcal{L}}{\partial \theta} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\theta}} \quad (14)$$

The metric tensor g_{ij} is a plane, so:

$$\begin{aligned} \bar{J} &= \frac{\partial \mathcal{L}}{\partial \dot{\theta}} = \text{constant} \\ &= m r^2 \dot{\theta} \end{aligned} \quad (15)$$

the conservation of angular momentum is:

$$\frac{d\bar{J}}{dt} = 0 \quad (16)$$

and the conservation of energy is:

$$\frac{dE}{dt} = 0 \quad (17)$$

The potential energy is therefore:

$$U = \frac{1}{2} m r^2 \dot{\theta}^2 = \frac{1}{2} \frac{\bar{J}^2}{m r^2} \quad (18)$$

and

$$F = \frac{\bar{J}^2}{2 m r^3} \quad (19)$$

1) 124(11) : Force Law and orbit due to Constant Torque

The Euler Lagrange equation is :

$$\frac{\partial L}{\partial r} = \frac{d}{dt} \frac{\partial L}{\partial \dot{r}} \quad (1)$$

w/ $L = \frac{1}{2} m (\dot{r}^2 - r^2 \dot{\theta}^2) - U(r)$ (2)

So: $F = m \ddot{r} = - \frac{\partial U}{\partial r}$ (3)

w/ $W_{12} = \int_1^2 \underline{F} \cdot d\underline{r}$ (4)

and $\underline{F} = - \nabla U$ (5)

Therefore: $\int_1^2 \underline{F} \cdot d\underline{r} = U_1 - U_2$ (6)

The work done is the change of potential energy:

$$W_{12} = U_2 - U_1 \quad (7)$$

Work is done by the star if $U_1 > U_2$. The

force is: $\int_1^2 \underline{F} \cdot d\underline{r} = - \int_1^2 dU = U_1 - U_2$ (8)

2) If F is positive valued the force is repulsive by convention. Therefore a repulsive force means that U_1 is greater than U_2 by convention. If F is negative valued the force is attractive by convention, so U_2 is greater than U_1 by convention. The initial state of potential is closer such that:

$$U_1 = 0 \quad \text{--- (9)}$$

and the final state is:

$$U_2 = \frac{1}{2} m r^2 \dot{\theta}^2 \quad \text{--- (10)}$$

Therefore:

$$U_2 > U_1 \quad \text{--- (11)}$$

and:

$$W_{12} > 0 \quad \text{--- (12)}$$

This means that work is done on a star by the force of spacetime. By convention, the force is negative valued, meaning that the star is attracted by the whirling spacetime.

Therefore:

$$F = m r \ddot{r} = - \frac{dU}{dr} = - m r \dot{\theta}^2 \quad \text{--- (13)}$$

Now make the change:

$$u = \frac{1}{r} \quad \text{--- (14)}$$

so

$$\frac{du}{d\theta} = - \frac{1}{r^2} \frac{dr}{d\theta} = - \frac{1}{r^2} \frac{dr}{dt} \frac{dt}{d\theta} = - \frac{1}{r^2} \frac{\dot{r}}{\dot{\theta}} \quad \text{--- (15)}$$

3) So $\frac{du}{d\theta} = -\frac{m}{J} \dot{r} \quad (16)$

and $\frac{d^2u}{d\theta^2} = -\frac{m^2}{J^2} r^2 \ddot{r} \quad (17)$

So $F(r) = -\frac{J^2}{m r^3} \frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) \quad (18)$

If the orbit is a logarithmic spiral:

$r = r_0 \exp(d\theta) \quad (19)$

and $\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) = \frac{d^2}{r} \quad (20)$

So $F(r) = -\frac{d^2 J^2}{m r^3} \quad (21)$

Re potential is $U(r) = \frac{d^2 J^2}{m r^2} \quad (22)$

so $d = 1/\sqrt{J} \quad (23)$

The velocity $v = \dot{r}^2 + r^2 \dot{\theta}^2$

$v^2 = \left(\frac{dJ}{mr} \right)^2 + \left(\frac{J}{mr} \right)^2 \quad (24)$

$v \rightarrow \infty \quad Bre^{-d\theta} = \frac{J}{m r_0}$

1) 124(12): Force Law for a Logarithmic Spiral
 (Maia and Thoma page 250)

In this development consider a star moving with velocity v in a plane. Let work be done on the star by a potential:

$$U(r) = -\frac{J^2 (1+d^2)}{2m} \frac{1}{r^2} \quad - (1)$$

so: $F(r) = -\frac{J^2 (1+d^2)}{mr^3} \quad - (2)$

The orbit is the logarithmic spiral:

$$r = r_0 \exp(d\theta) \quad - (3)$$

Eq. (1) is the potential energy due to constant specific rotation.

We have:

$$\theta(t) = \frac{1}{2d} \log_e \left(\frac{2dJt}{mr_0^2} + c \right) \quad - (4)$$

$$r(t) = \left(\frac{2dJt}{m} + r_0^2 c \right)^{1/2} \quad - (5)$$

Animate eqns. (4) and (5).

Also: $\dot{\theta} = \frac{J}{mr^2} \quad - (6)$

$$\dot{r} = \frac{dJ}{mr} \quad - (7)$$

2)

$$v^2 = r \cdot \ddot{\alpha} + r^2 \dot{\theta}^2 \quad (8)$$

In the context of the quantum emission phenomenon can further transform into a quantum potential of the system by Bohm's from an energy aspect, the quantum potential exists in an energy context, so that distant elements participating in it will exist together as if they were co-located or partially co-located. The energy distribution aspects of the quantum potential and the vacuum energy amplification/distribution aspects of the quantum potential are revolutionary.

So:

$$J = m v r (1 + d^2)^{-1/2} \quad (9)$$

Here J , d , r_0 and C are constants. Eq. (9) shows that the orbit evolves to that of a

circle of radius $(1 + d^2)^{-1/2} r$. In a circular orbit both v and r are constant.

Let: $R = (1 + d^2)^{-1/2} r \quad (10)$

and so:

$$J = m v R = \text{constant} \quad (11)$$

The basic equation is $m \ddot{r} = - \frac{m c^2}{r^2} \quad (7.21)$

$$\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} = - \frac{m c^2}{J^2} F(r) \quad (12)$$

124(13): Resonance & Inverse Square Law of Gravitational

Attraction

Start with the equation:

$$\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} = -\frac{mr^2}{J^2} F(r) \quad \text{--- (1)}$$

Kepler's laws and Newtonian dynamics are given by:

$$F(r) = -\frac{k}{r^2} \quad \text{--- (2)}$$

which is the inverse square law of gravitational attraction.

Therefore:

$$\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} = \left(\frac{mk}{J^2} \right) \quad \text{--- (3)}$$

Now introduce an oscillatory structure into the familiar

inverse square law:

$$F(r) = -\frac{k}{r^2} \cos(d\theta) \quad \text{--- (4)}$$

Then:

$$\boxed{\frac{d^2}{d\theta^2} \left(\frac{1}{r} \right) + \frac{1}{r} = \left(\frac{mk}{J^2} \right) \cos(d\theta)} \quad \text{--- (5)}$$

This is a Bessel's: Euler resonance equation.

At resonance:

2)

$$\frac{1}{r} = \frac{mk \cos(\alpha\theta)}{\sqrt{2(d^2 - 1)^{1/2}}} \quad (6)$$

so if: $d^2 = 1, \alpha = \pm 1$ — (7)

then $\frac{1}{r} \rightarrow \infty, r \rightarrow 0$ — (8)

and the force $F(r)$ becomes negative i.e. $\frac{1}{r^2}$ and the mass m is attracted to the mass M by an infinite force, and the system implodes.

Application to Counter Gravitation

When: $\cos(\alpha\theta) < 0$ — (9)

Then: $F(r) > 0$ — (10)

and the force becomes repulsive. At the resonance point (i) the object m is repelled from M . This is an example of resonant counter-gravitation, where the system explodes.