

108(a) : New Cubic Equation for the Relativistic Kepler Problem.

Consider again the potential energy of the system:

$$V = mc^2 \left(\frac{1}{2} - \frac{r_s}{r} \right) + \frac{nL^2}{2r^2} - nL^2 \frac{r_s}{r^3} \quad (1)$$

In the standard model r_s is a constant but in ECE it may be regarded as a variable, so:

$$\frac{\partial V}{\partial r_s} = -\frac{mc^2}{r} - \frac{nL^2}{r^3} \quad (2)$$

Therefore:

$$\left(\frac{\partial V}{\partial r_s} \right) r^3 + mc^2 r^2 + nL^2 = 0 \quad (3)$$

This is a cubic equation with three roots, in which L is a constant and where $(\partial V / \partial r_s)$ can be regarded as an adjustable parameter. In

ECE: $|r_s| = T/R \quad (4)$

The rate of change of V with r is:

$$\frac{\partial V}{\partial r} = \frac{m}{r^2} \left(c^2 r_s - \frac{L^2}{r} + 3L^2 \frac{r_s}{r^2} \right) \quad (4)$$

Now use:

2)

$$\frac{\partial \bar{V}}{\partial r_s} = \frac{\partial V}{\partial r} \frac{\partial r}{\partial r_s} \quad - (5)$$

So:

$$\frac{\partial r}{\partial r_s} = \left(\frac{\partial V}{\partial r_s} \right) / \left(\frac{\partial V}{\partial r} \right) \quad - (6)$$

$$\frac{\partial r}{\partial r_s} = - \left(\frac{mc^2}{r} + \frac{nL^2}{r^3} \right) \left(\frac{m}{r^2} \left(c^2 r_s - \frac{L^2}{r} + 3 L^2 \frac{r_s}{r^2} \right) \right)^{-1}$$

- (7)

Thus:

$$\frac{\partial r}{\partial r_s} = - \left(rc^2 + \frac{L^2}{r} \right) \left(c^2 r_s - \frac{L^2}{r} + 3 L^2 \frac{r_s}{r^2} \right)^{-1}$$

and

$$\frac{\partial r_s}{\partial r} = - \left(c^2 r_s - \frac{L^2}{r} + 3 L^2 \frac{r_s}{r^2} \right) / \left(rc^2 + \frac{L^2}{r} \right)$$

$$\frac{\partial r_s}{\partial r} = r \left(\frac{L^2}{r} - c^2 r_s - 3 L^2 \frac{r_s}{r^2} \right) / \left(r^2 c^2 + L^2 \right)$$

$$r_s = \int \frac{r \left(\frac{L^2}{r} - c^2 r_s - 3 L^2 \frac{r_s}{r^2} \right) dr}{(r^2 c^2 + L^2)}$$

3) Now we:

$$L = r^2 \frac{d\phi}{dr} \quad - (9)$$

and:

$$r = \int \left(rc^2 + \frac{L^2}{r} \right) \left(\frac{L^2}{r} - c^2 r_s - 3 \frac{L^2 r_s}{r^2} \right)^{-1} dr_s \quad - (10)$$

Thus r_s may be integrated out from eq. (10) and r graphed as a function of L . For example

if $|r_s| = T/R \quad - (11)$

then for a given $(T/R)_0$: $- (12)$

$$r = \int_0^{(T/R)_0} \left(rc^2 + \frac{L^2}{r} \right) \left(\frac{L^2}{r} - c^2 r_s - 3 \frac{L^2 r_s}{r^2} \right)^{-1} dr_s$$

This gives an equation for r in terms of $(T/R)_0$ and L .

As $r_s \rightarrow 0$ in eq. (7):

$$\frac{dr}{dL} \rightarrow \frac{r}{L^2} \left(rc^2 + \frac{L^2}{r} \right) \quad - (13)$$

4)

Therefore as $r_s \rightarrow 0$:

$$r \rightarrow \left(1 + \left(\frac{rc}{L} \right)^2 \right) r_s \quad - (14)$$

This gives the quadratic equation:

$$\left(\frac{c^2 r_s}{L^2} \right) r^2 - r + r_s = 0 \quad - (15)$$

$$\text{So } r = \frac{2L^2}{c^2 r_s} \left(1 \pm \left(1 - 4 \left(\frac{c r_s}{L} \right)^2 \right)^{1/2} \right)$$

$$\rightarrow \frac{2L^2}{c^2 r_s} = \frac{2r^4}{c^2 r_s} \left(\frac{d\phi}{d\tau} \right)^2$$

$$\text{So: } r^3 \rightarrow \frac{1}{2} c^2 r_s \left(\frac{d\phi}{d\tau} \right)^{-2} \quad - (16)$$

It is known that L is constant:

$$L = r^2 \frac{d\phi}{d\tau} = \text{constant} \quad - (17)$$

$$\text{So as } r \rightarrow 0, \frac{d\phi}{d\tau} \rightarrow \infty \quad - (18)$$

From (17) and (18):

$$r \rightarrow 0 \text{ as } r_s \rightarrow 0 \quad - (19)$$