

Notes 106(3) : Orbits of Binary Pulsars

The first binary pulsar was found in 1974 by Hulse and Taylor. In 2004 a second one was found at Jodrell bank. The HT pulsar consists of a pulsar (a neutron star) with a pulsation period of 39 milliseconds and a second neutron star in an elongated orbit of period 7.75 hours. Each neutron star is 1.4 million times the mass of the sun. The orbit is gradually shrinking by about 3.1 mm per orbit. In 300 million years the two stars will collide. The orbital precession is 4.2 degrees a year. In ECE theory a constant orbital precession is given by a constant T/R :

$$-\frac{T}{R} = \frac{2MG}{c^2} := r_s \quad - (1)$$

However, the orbit of the binary pulsar is a spiral inward towards each other, so $-T/R$ is not constant. It is time dependent. This means that the energy:

$$V(r) = \frac{mc^2}{2} \left(-\frac{r_s}{r} + \frac{a^2}{r^2} - \frac{r_s a^2}{r^3} \right) \quad - (2)$$

is getting smaller with time.

Here:

$$\frac{mc^2}{2} a^2 = \frac{L^2}{2m} \quad - (3)$$

$$\frac{mc^2}{2} a^2 r_s = \frac{GM L^2}{c^2 m} \quad - (4)$$

2) In the standard model the loss of potential energy is postulated to be gravitational radiation. For several reasons this postulate is incorrect. Therefore we replace it by a modelling of T/R , which is given more fully in papers 105 and notes 106(1) and 106(2).

In ECE theory the advance of the perihelion is given by:

$$\delta\phi = -\frac{3\pi}{A(1-e^2)} \left(\frac{T}{R}\right) \quad \text{--- (5)}$$

which is calculated from an angular frequency:

$$\omega_r^2 = \frac{1}{m} \left(\frac{d^2V}{dr^2} \right)_{r=r_{\text{outer}}}$$

$$\omega_r^2 = \omega_\phi^2 \left(1 - \frac{3r_s^2}{a^2} \right)^{1/2} \quad \text{--- (6)}$$

with r_s given by eq. (1).

Computer graphics

1) From eq. (6) graph ω_r^2 against t with R/T modelled as a function of time. As the two stars spiral inward, ω_r^2 will increase. Fit to experimental data of -3.1 mm per orbit.

3) 2) The equation of motion of a mass m orbiting a mass M is:

$$\bar{V} = \frac{1}{2} m \left(\frac{dr}{dt} \right)^2 = \left(\frac{E^2}{2mc^2} - \frac{1}{2} mc^2 \right) + \frac{GMm}{r} - \frac{L^2}{2mr^2} - 2 \left(\frac{T}{R} \right) \frac{L^2}{mr^3} \quad (7)$$

Graph \bar{V} as a function of time dependent T/R .

A time dependent V/R arises from a time dependence of the spin constant, i.e. the simplest case:

$$\omega_r = \frac{T}{R} \quad (8)$$

For universal gravitation this is constant, however in the case of relativity it is time dependent, i.e. Newton's laws are slowly changing with evolution (Pioneer anomaly).